

VI LAPIS 2018: La Plata International School on Astronomy and Geophysics
"COSMOLOGY IN THE ERA OF LARGE SURVEYS"
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LAB I: CAMB

<https://camb.info/>

How to run camb?

```
> make clean  
> make  
> ./camb file.ini (params.ini)
```

1) Choose a cosmological model (namely fiducial model) and plot the Cl's (TT,EE,TE,BB). Compare them with the Cl's of other model such that:

- a) the baryon density is 20% higher than the one of the fiducial model;
- b) the CDM density is 15% lower than the one of the fiducial model;
- c) the spectral index is 10% higher than the one of the fiducial model;
- d) the optical depth is 25% lower than the one of the fiducial model.

For all cases, plot also the relative differences:

$\Delta C_l = [C_l - C_l(\text{fiducial model})]/C_l(\text{fiducial model})$

2) Modify the power spectrum in CAMB, and plot the new P(k) and Cl's.

- a) $P(k) = A \cdot k^{(n-1)} \cos(k)$
- b) $P(k) = A \cdot k^{(n-1)} \cos(b \cdot k)$, where b is a new parameter of the model.

3) Compute the Cl's for a cosmological model with dynamical dark energy using an equation of state of the form $w(a) = w_0 + w_a \cdot (1-a)$, with $w_0 = -0.92 \pm 0.11$ and $w_a = -0.32 \pm 0.40$ (ref.: Sanchez et. al 2017).